"You miss $100 \%$ of the shots you never take." Wayne Gretzky

## Notices

HW2 is due Friday.
Late HW1 papers?
HW1 passed back on Wednesday, be sure to sit in the row you put on your HW. HW1 solutions posted after class on Wednesday.

## Jupiter

## Jupiter's Structure

- Top: H clouds over
- Liquid H over
- Liquid metallic H over
- Core Rocky, metalic core.

Jovian Planet Structure (our 2 ${ }^{\text {nd }}$ of 5!)


## (Warm interior) Icy moon structure

Crust made mostly of water-ice.
Underneath lies a vast liquid layer of water Rocky mantle. Rocky (iron) core.

Densities ~ 2 g/cc depending on how much rock to water/ice.

Structure \#3

## (Cold interior) Icy moon structure

Rock/ice mixture surface/mantle Rocky (iron-mix) core.

Densities ~ 2 g/cc depending on how mucb rock to water/ice.

## Structure \#4

## Densities (in g/cc):

Rocky worlds: 3.5 - 5.5 (Terrestrial planets $\sim 5$ )

Gas Giants: $0.7-1.5$

Icy moons: ~2-3


## Jupiter has rings!

## Jupiter's Rings

The shape of the rings is determined by small (shepherd) moons, Jupiter's strong magnetic field, and of course Jupiter's gravity.

The particles are extremely smallmicrometer sized!

The particles in the rings only last for about 1,000 years before falling into Jupiter. But they are continually replenished by impacts on Jupiter's moons.


## Jupiter's moons

Jupiter has at least 63 moons. The 4 largest are known as the Galilean moons, named after their discoverer.

Io is slightly larger than our Moon while Callisto is about the same size as Mercury. Ganymede is $10 \%$ larger than Mercury!

All of these moons are locked in synchronous rotation. Only one side ever faces Jupiter.


## Io, the innermost Galilean moon




## Io, the innermost Galilean moon

*Io is the most volcanically active body in the solar system! * There are no noticeable impact craters on Io. *The heating is generated by tidal forces. Io's orbit is elliptical because of a gravitational tug-of-war between Jupiter and Europa, the next large moon.
This stretching and pulling heats Io's interior.

## Io is covered with active volcanoes!



## Io's structure: Terrestrial

Io has 3 layers: a thin, brittle rocky crust, a thick, molten-rock mantle, and a solid, core made mostly of iron.

Density=
$3.5 \mathrm{~g} / \mathrm{cc}$

## Io's atmosphere

Io has a thin, $\mathrm{SO}_{2}$ atmosphere. Some freezes out on the night side and evaporates on the day side.

## Europa




## Europa

Also caught in a tug-of-war between massive Jupiter and other large moons, Europa has a fairly young surface, with only a few impact craters.

## Europa's structure: Icy moon, warm interior.

Europa has a thin (3 miles thick) crust made mostly of water-ice. Underneath lies a vast liquid layer of water, perhaps 30 miles thick.
But most of the moon is made of rock, with perhaps an iron core.

Density = $3.0 \mathrm{~g} / \mathrm{cc}$
Europa has a thin, Oxygen atmosphere created by charged particles hitting the surface.
Comparable to
Earth's atmosphere at an altitude of over 100 miles!

## Europa

has a vast undercrust ocean. Kept liquid by volcanic vents.

## Ganymede



## Ganymede

As we move out from Jupiter, the tidal flexing becomes less, and the surfaces get colder and older. Ganymede is the largest moon in the solar system and also generates it's own magnetic field, just like the Earth!


## Ganymede's interior

Like Europa, Ganymede has a thin, mostly ice crust. Underneath is a thick, water-ice layer (mostly frozen): Under that is a rocky mantle surrounding an iron core:

Density $=1.9 \mathrm{~g} / \mathrm{cc}$

Like Europa, Ganymede has an extremely thin Oxygen atmosphere created by charged particles hitting the ice surface.

## Callisto

## Callisto's surface is old and heavily cratered.



## Callisto's structure

Being the furthest moon from Jupiter and under the least tidal stress, Callisto did not have time to differentiate much before it became frozen.

## Density = $1.9 \mathrm{~g} / \mathrm{cc}$

However, Callisto shows signs of a variable magnetic field, an indication of a subterranean, salty ocean.

## Older \& colder the farther away from Jupiter.



Io < few million, Europa ~20-180 million years, Ganymede $\sim 3$ billion years, Callisto: 4-4.5 billion years old

## Structure changes as they get further from Jupiter.

# Surface Age: <br> Io < few million 

## Europa ~20-180 million years

Density $=3.5 \mathrm{~g} / \mathrm{cc} \quad$ Density $=3.0 \mathrm{~g} / \mathrm{cc}$
Ganymede ~3 billion years


Callisto: 4-4.5
billion years old

Density $=1.9 \mathrm{~g} / \mathrm{cc}$
Density $=1.8 \mathrm{~g} / \mathrm{cc}$

## Structure changes slightly as they get further from Jupiter.

Notice how
prevalent water is in the solar system!
Europa, Callisto (liquid?)
Ganymede,
Moon, Mercury
(ice)

## Jupiter also has lots of smaller, irregular moons. Jupiter has 63 moons in total.




## Saturn

## Infrared <br> view

# Saturn's structure is very similar to Jupiter's. Clouds which thicken to liquid Metallic hydrogen interior Rocky core Density $=0.7$ less than water! 



Solar day $=10$ hours 40 minutes
This spin is so fast that Saturn is visibly larger at the equator then the poles. 95 times more massive then Earth.


## Saturn's Satellites and Ring Structure 56 known moons



This graphic is available in color if required.

## Describe <br> Saturn's rings.

## Saturn's Rings.

* Are extremely thin: 200,000km across, but typically about 200m thick!
* Made of millions of particles- but not very thick!
-1 cm to 10 m in diameter
-mostly ice or ice-covered rocks
- all the ring material, compressed into a rock would only be 100 km across.



## Cassini Visual and Infrared Mapping Spectrometer







Shepherd moons shape the rings. But do they make them stable? An open question. Prior to Galileo's study of the rings, it was thought they were about 200Myrs old and would only last another

200Myrs.

