

**"Ah, but a man's reach should exceed  
his grasp, or what's a heaven for?"  
Robert Browning**

**Group Project 1 is due. I will collect them at the  
beginning of class.**

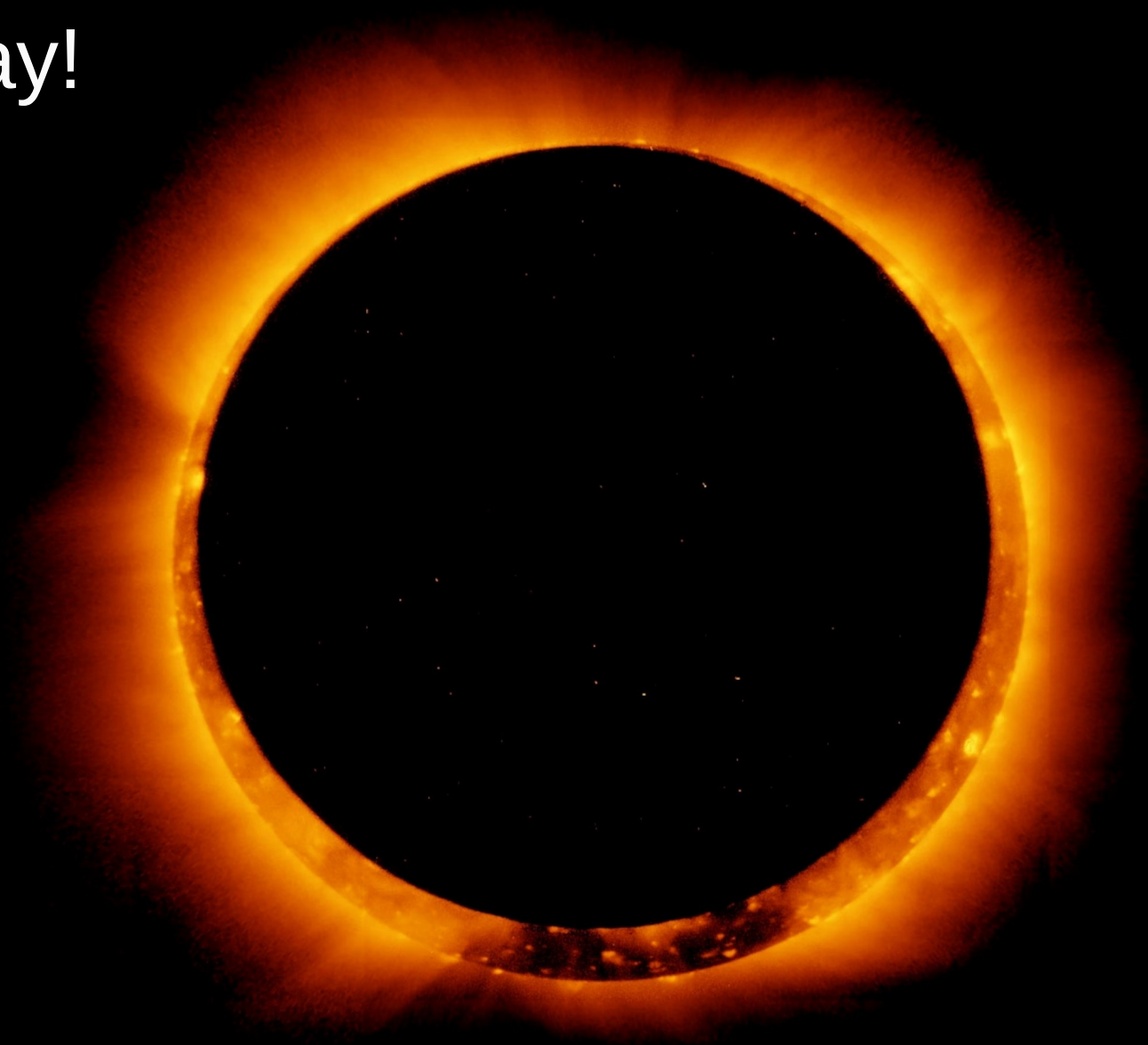
Test 2 on Wednesday in class.

Sample Tests and study guide on the course web page now.

I will be gone April 4-11. 3 classes will be put on the course web page as videos with powerpoints.

I will also put a quiz or 2 on blackboard and a discussion page (for points) on blackboard.

Eclipse 1 week  
from today!



Partial: be sure to always wear eye protection



“The clouds I can handle, but I can’t fight with an eclipse” Stephanie Meyer, Eclipse





Protostars: gravity

Main Sequence:  
fusion  $H \rightarrow He$  in  
their cores

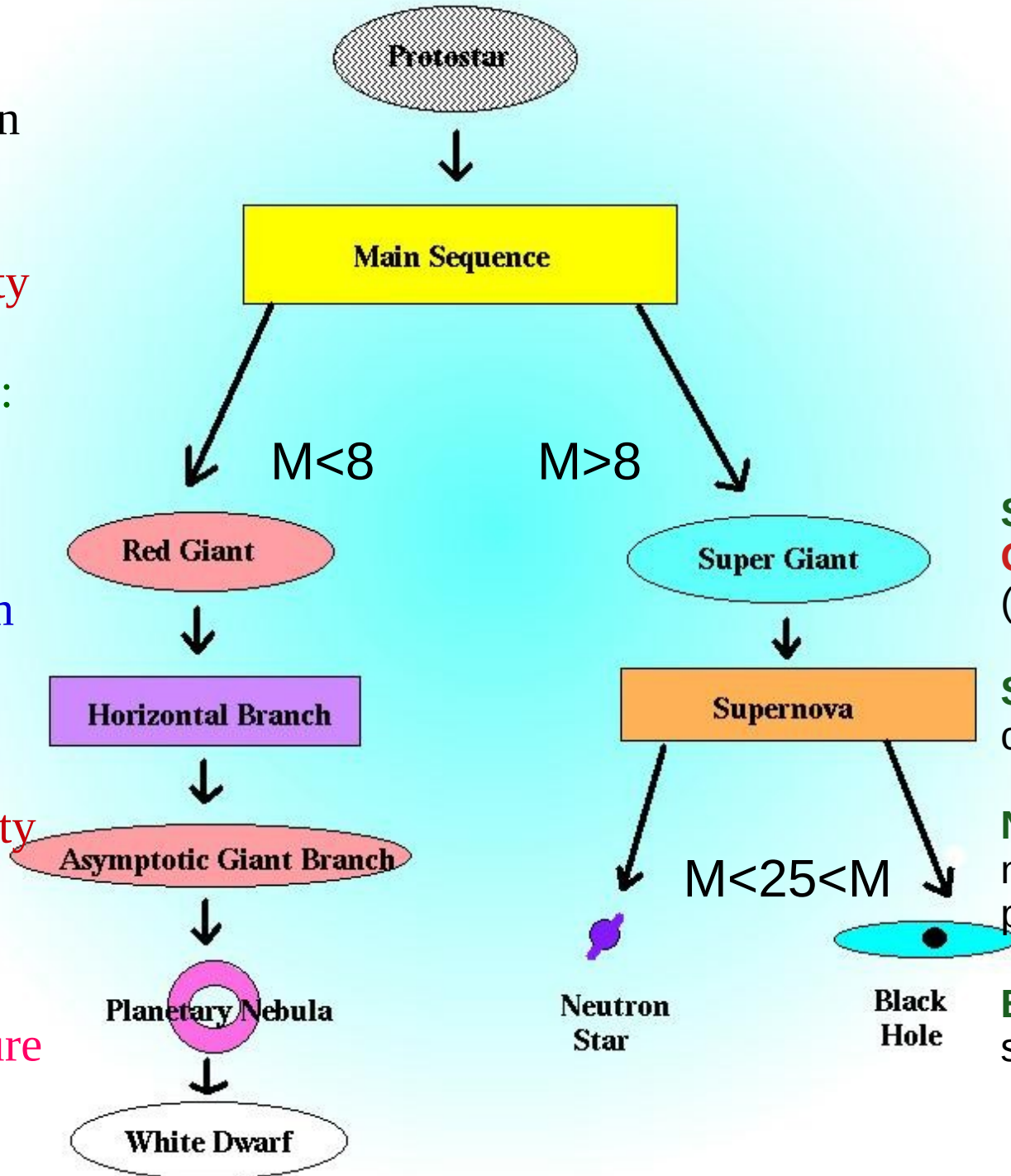
Red giants: gravity

Horizontal branch:  
fusion of  
 $He \rightarrow C$

AGB: energy from  
gravity

Planetary nebula:  
energy from gravity

White dwarf:  
electron  
degeneracy pressure



**Supergiant**:  
Gravity + fusion  
(He, C, N, O, Si)

**Supernova**:  
collapse, no support.

**Neutron Star**:  
neutron degeneracy  
pressure.

**Black hole**: no  
support.

# End States of Stars

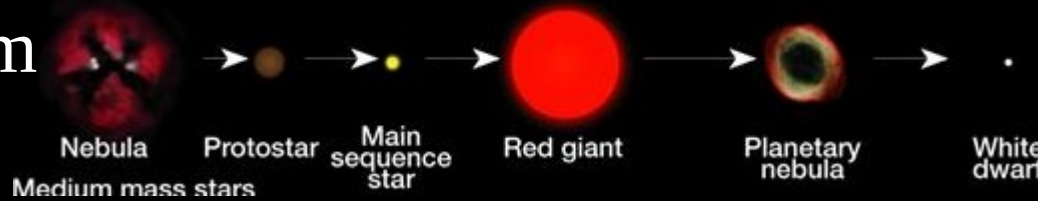
White dwarfs: electron degeneracy pressure. Earth-sized,  $0.6M_{\text{sun}}$

Neutron Stars: neutron degeneracy pressure. City-sized,  $1.4M_{\text{Sun}}$

Black holes: no support. Size is a point, but  $R_{\text{Sch}}=3M$  (mass in  $M_{\text{Sun}}$ ,  $R$  in km). Eventually black holes evaporate.

If we just keep going back to previous generations of stars, what happens to the 'metals'?

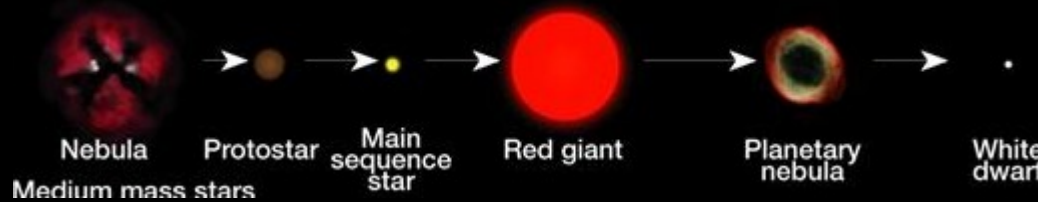
Our solar system  
2% metals



Planetary nebula  
returns 3% metals

Before our Sun?

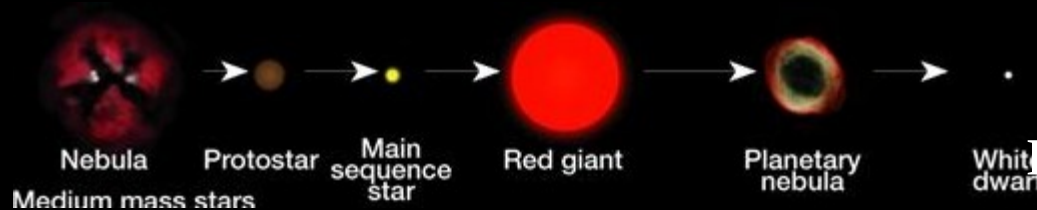
Previous star:  
1% metals



Planetary nebula  
returns 2% metals

And before that?

Previous star:  
0.5% metals



Planetary nebula  
returns 1% metals



# Stellar Populations

As stars recycle heavier material back into space, newly formed stars incorporate this material.

Thus newer stars have more “metals” than old stars.

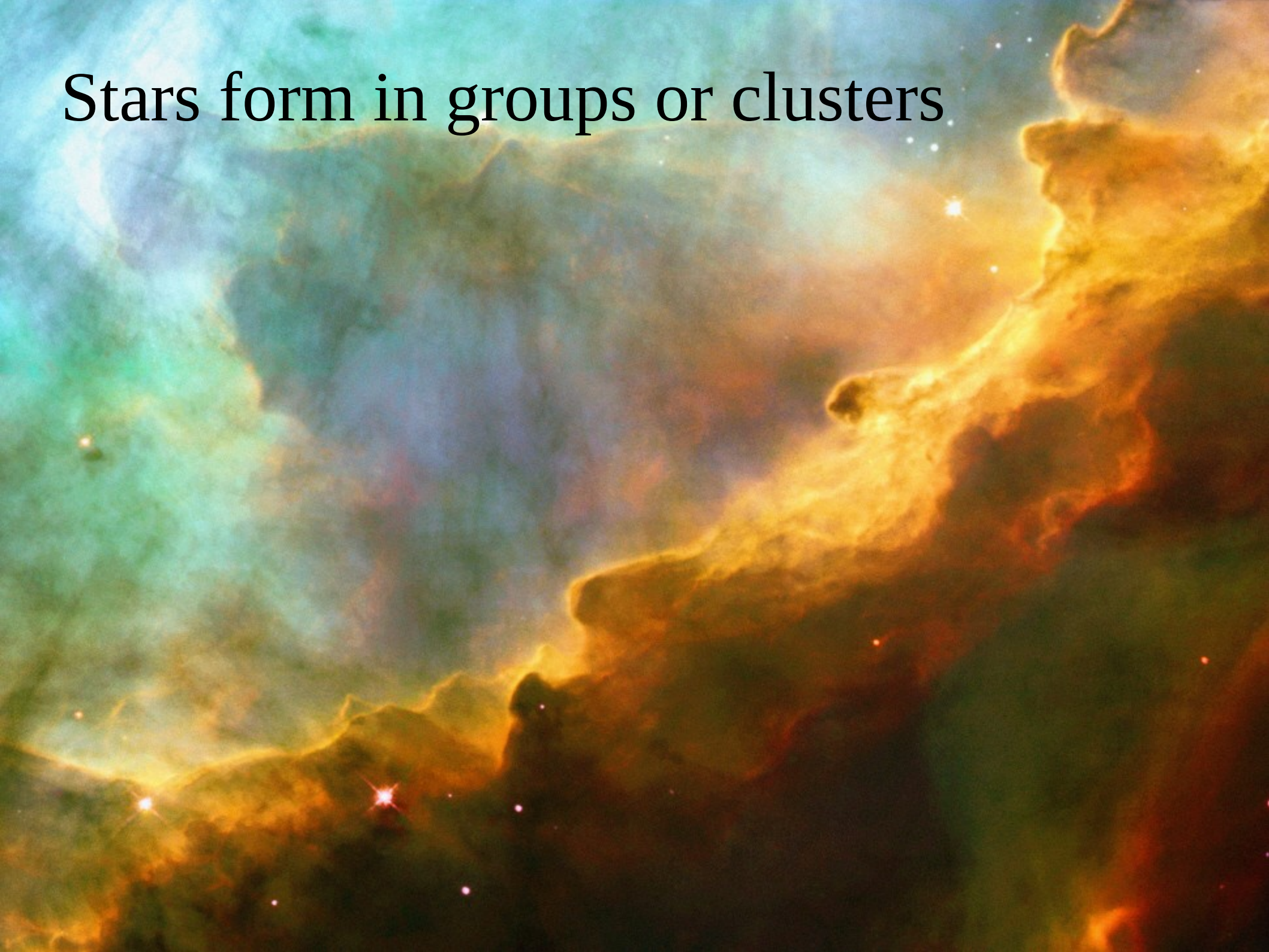
This gives us populations:

I: Stars like our Sun- they have the most metals  
(2%)

II: Metal-poor stars- formed in the galaxy before our Sun: have some, but little metals (<0.1%)

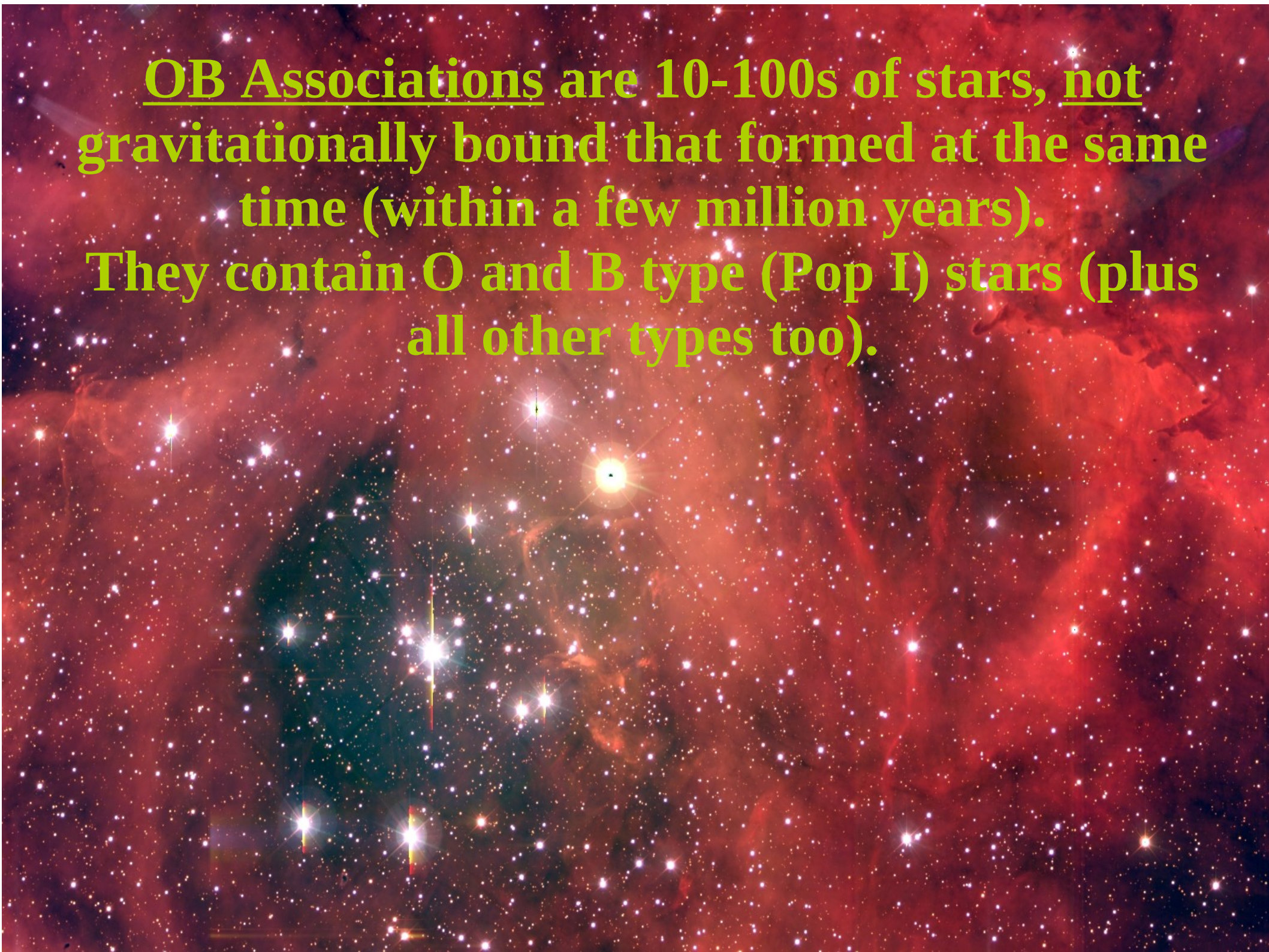
III: No metals at all: made only of H and He. The first stars.

Stars form in groups or clusters



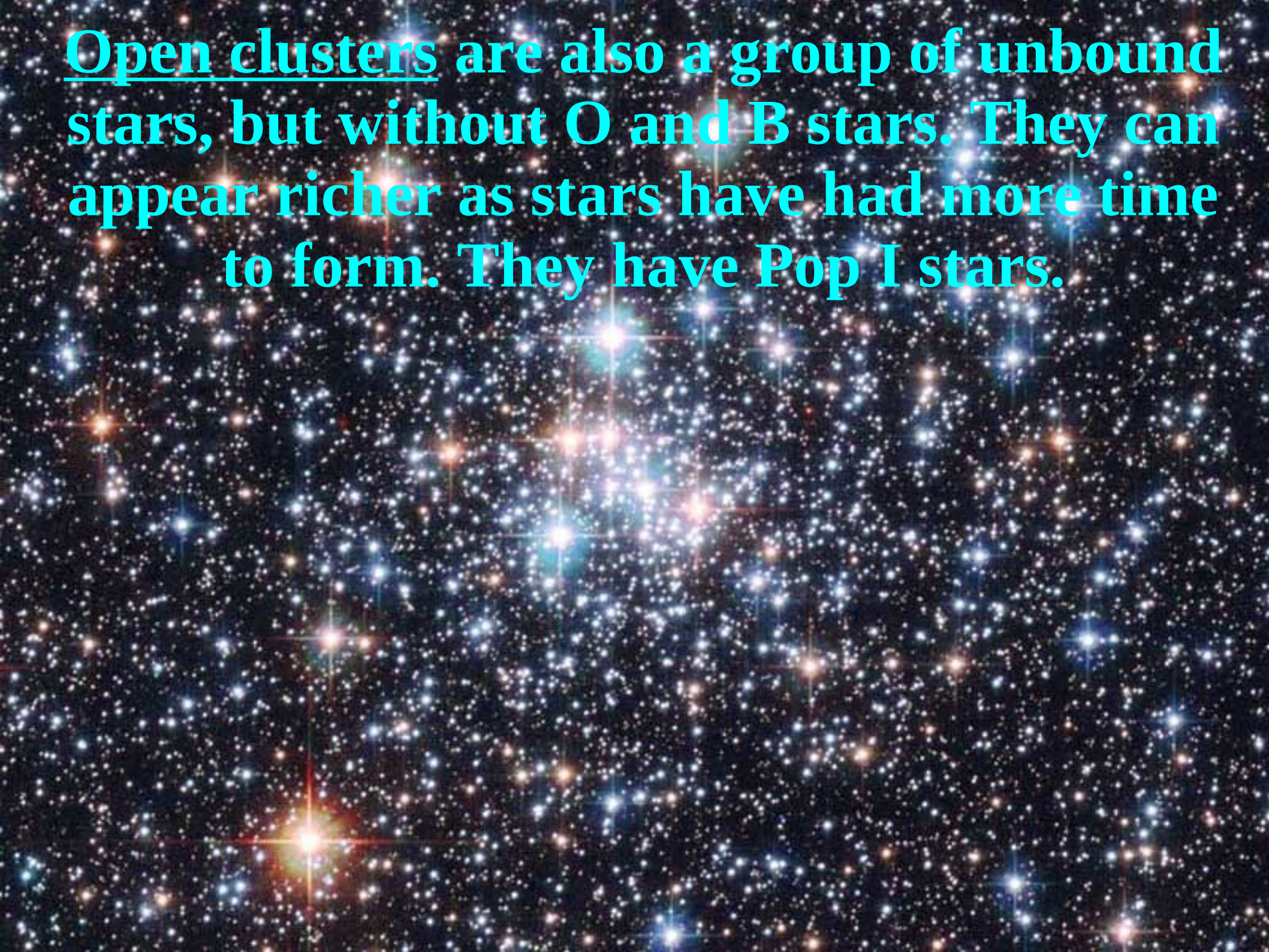


**OB Associations are 10-100s of stars, not gravitationally bound that formed at the same time (within a few million years). They contain O and B type (Pop I) stars (plus all other types too).**





Open clusters are also a group of unbound stars, but without O and B stars. They can appear richer as stars have had more time to form. They have Pop I stars.





Globular clusters contain up to millions of stars. They are gravitationally bound together and are very old, with only low mass stars remaining.

They have Population II stars



A dense field of stars, likely representing an open cluster, with a bright central region. The stars are of various colors and sizes, creating a rich, multi-colored appearance. The background is dark, making the individual points of light stand out.

**Our Sun was most likely formed  
within an open cluster that has  
since dissipated.**

**A nearby supernova caused our  
pre-solar cloud to collapse.**



# Take aways

Stars form in groups.

Most groups are not gravitationally bound, but the rich ones (globular clusters) are.

Supernova come in two types:  
Type I: Exploding white dwarfs  
Type II: Exploding massive stars.

**Why is this distinction important?** A Type Ia always has the same amount of mass in the explosion, a Type II can vary widely.

**How can you tell them apart?** Type II have H and He, Type I do not.

# Take aways

Supernova have 2 types.

Type Ia are exploding white dwarfs  
and are always the same brightness.

Great for determining distances!

# Quiz 10: Our Sun...

- A) is unusual as it is not in a binary or multiple star system.
- B) will eventually use up its fuel.
- C) will not explode.
- D) A, B, and C are all true.
- E) None of the above are true.

Q&A Time! What provides the support for white dwarf stars?

A) Fusion  $H \rightarrow He$

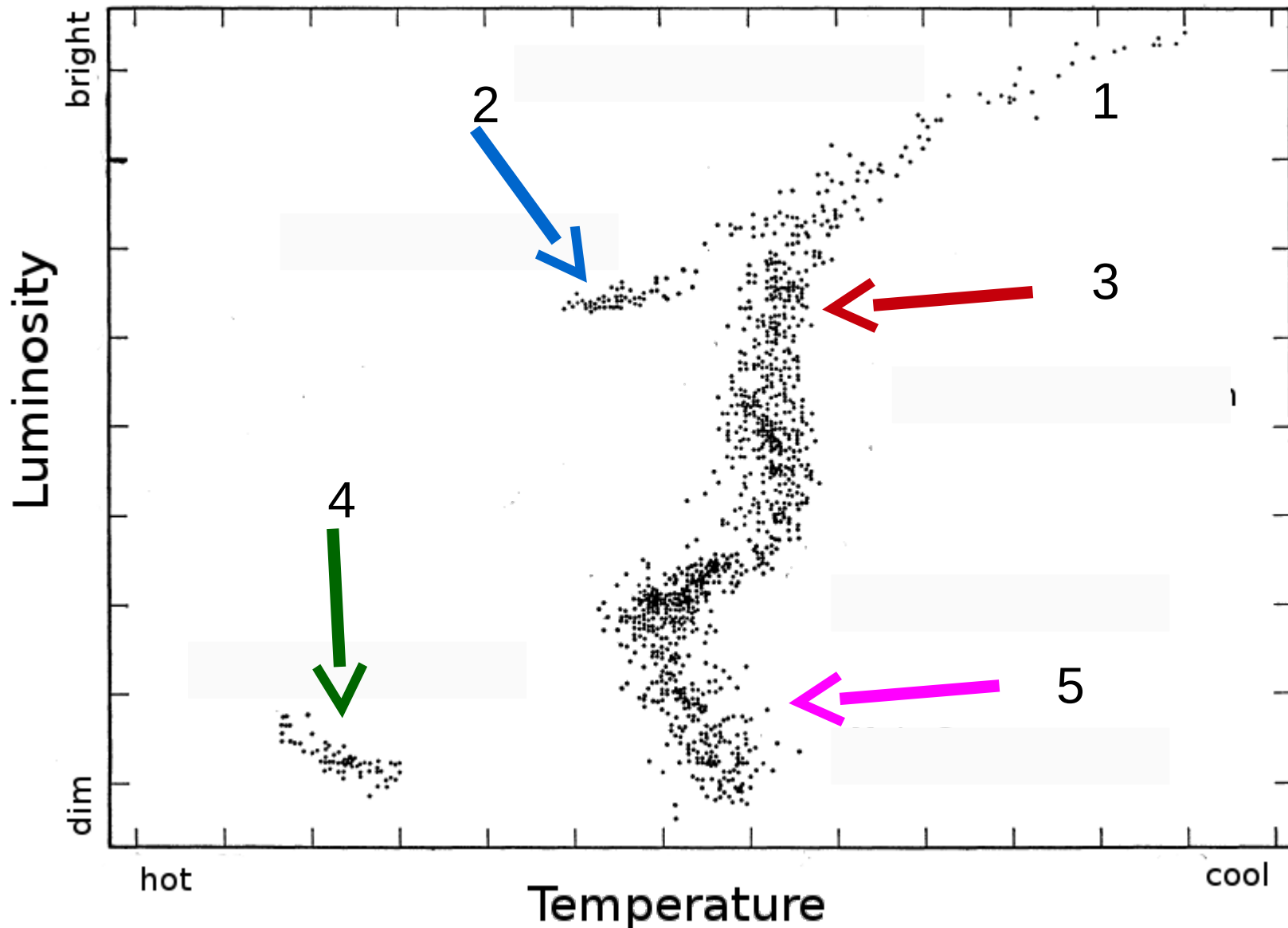
B) Fusion  $He \rightarrow C$

C) Gravity

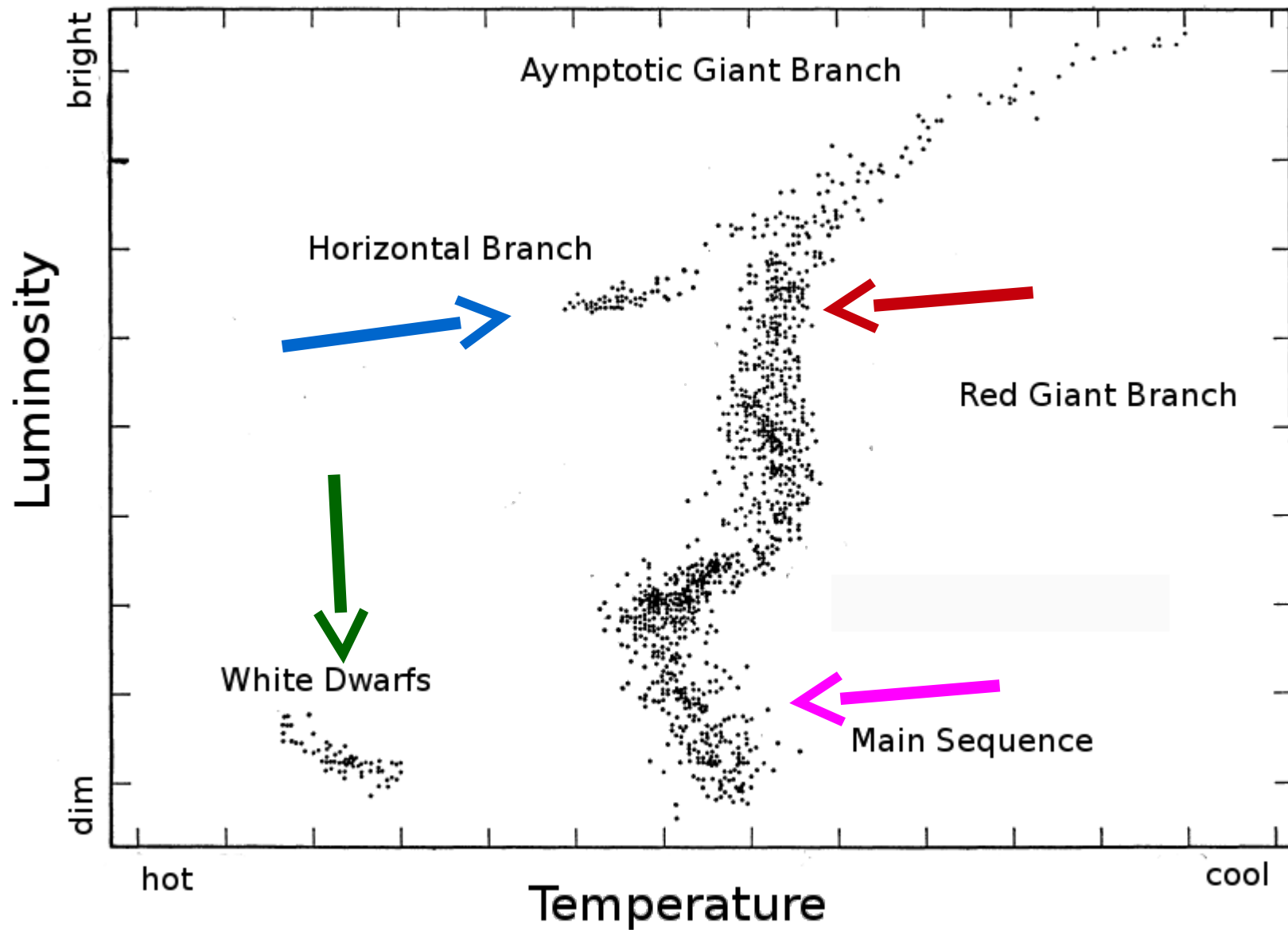
D) Electron degeneracy pressure

E) Neutron degeneracy pressure

# Q&A Time! What stages of evolution are the arrows pointing to?







Q&A Time! If at the same distance, which star should be brightest?

- A) Sun-like star
- B) White dwarf
- C) Supergiant
- D) Neutron star
- E) There is no way to tell

On to Galaxies....





What is a galaxy? What are galaxies made of?  
How big are galaxies?





# Galaxies

Galaxies are collections of stars and gas (and dust) that are gravitationally bound together.



# Galaxies

We will start with the easiest galaxy to study: Our own Milky Way.

Our Milky Way galaxy contains over 200 billion stars (that's  $2 \times 10^{11}$ !)