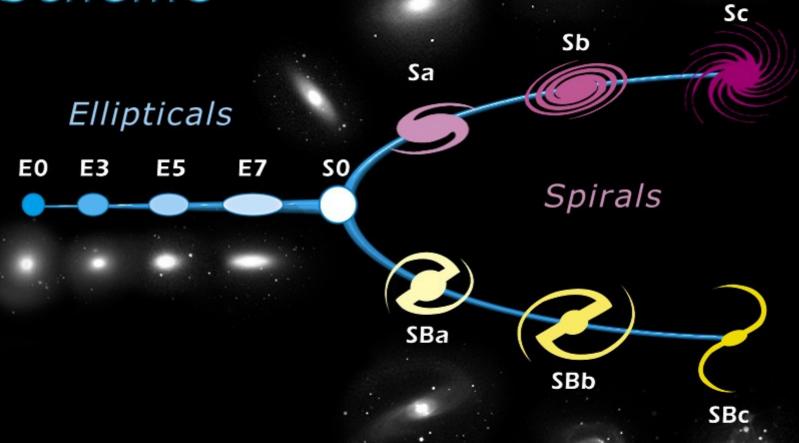


Edwin Hubble's Classification Scheme





Irregular galaxies:

These are not on the Hubble tuning fork. They do not have a disk, nor are they elliptical.

They tend to be young, gas-rich and contain Pop I stars.

Distances. How far away is everything and how can I tell?



We use something called the distance ladder.

That is, each step depends on the previous one.

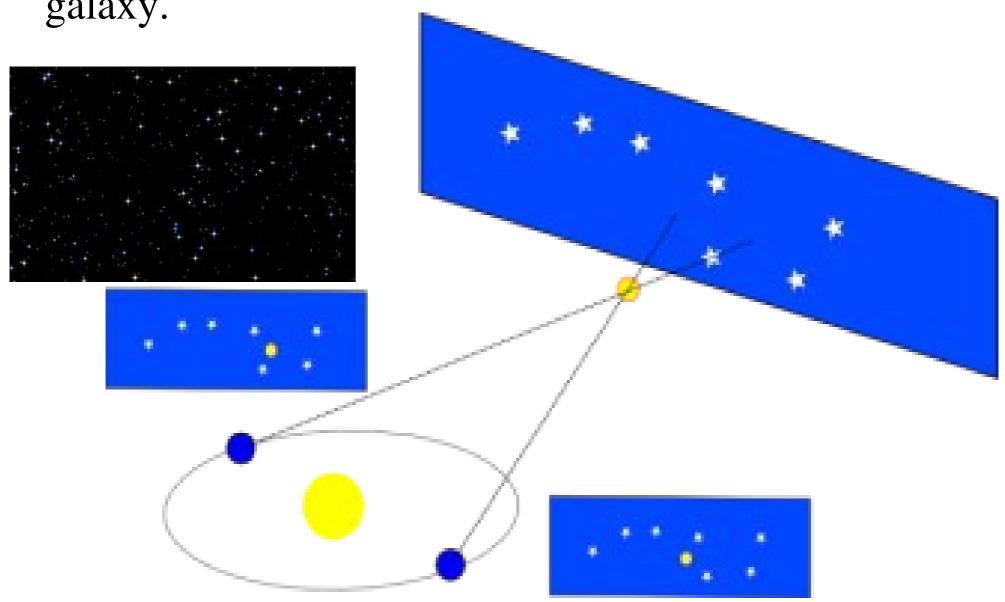


Distances.

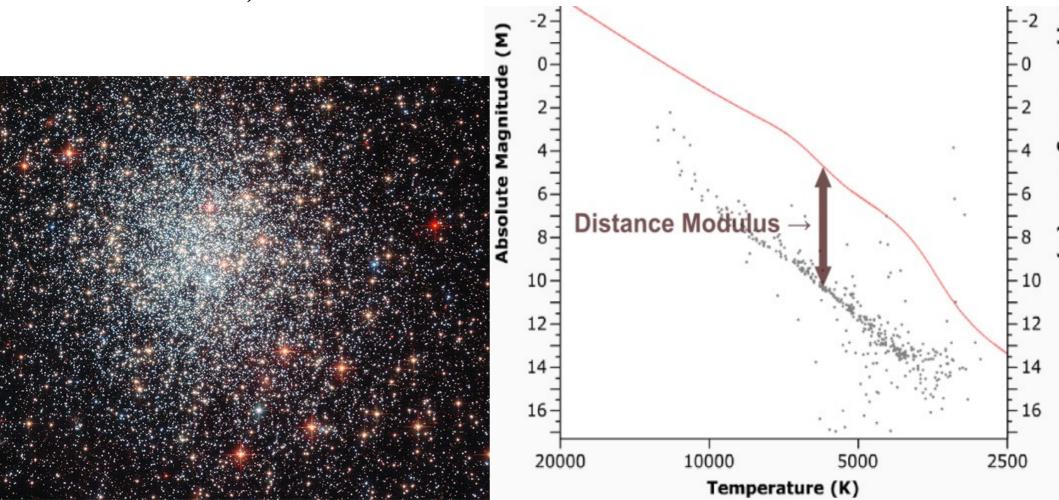
How far away is everything and how can I tell?

There will be summary slides

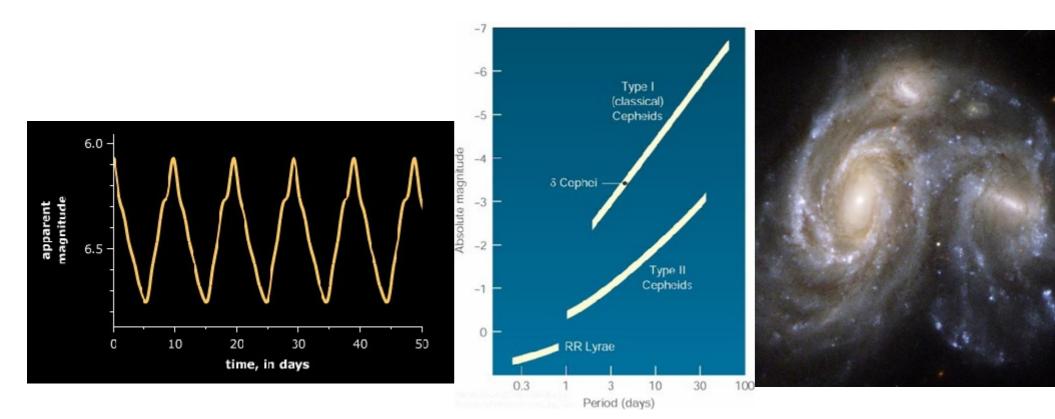
• Parallax- good for nearby stars. GAIA- half our galaxy.



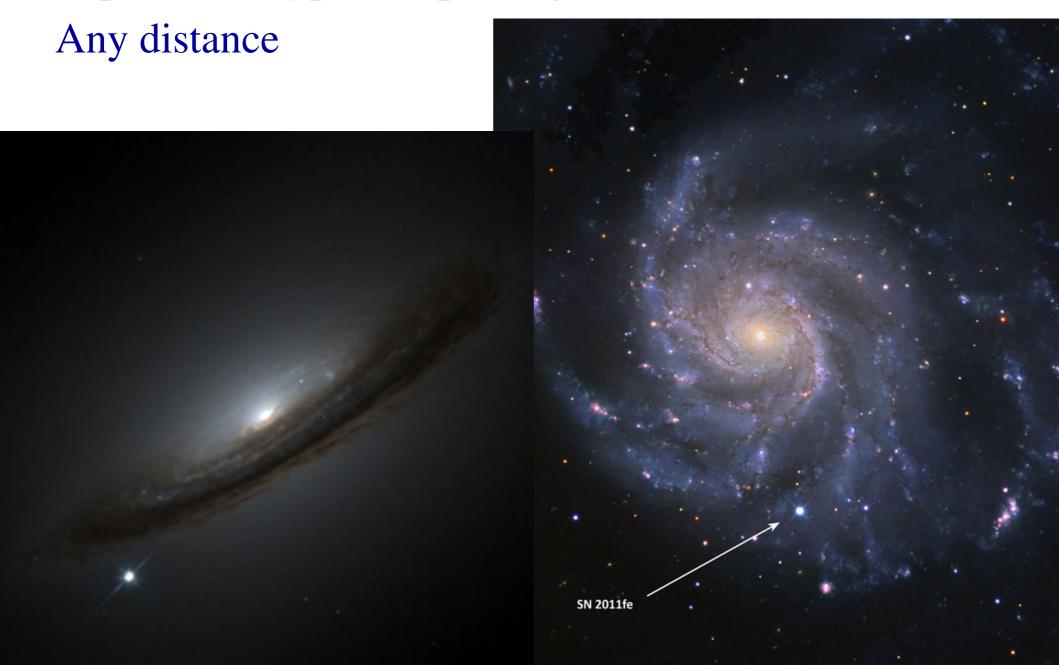
- Parallax- good for nearby stars
- Main sequence fitting- Good for nearby galaxies where individual stars can be resolved (true for eclipsing binaries too).

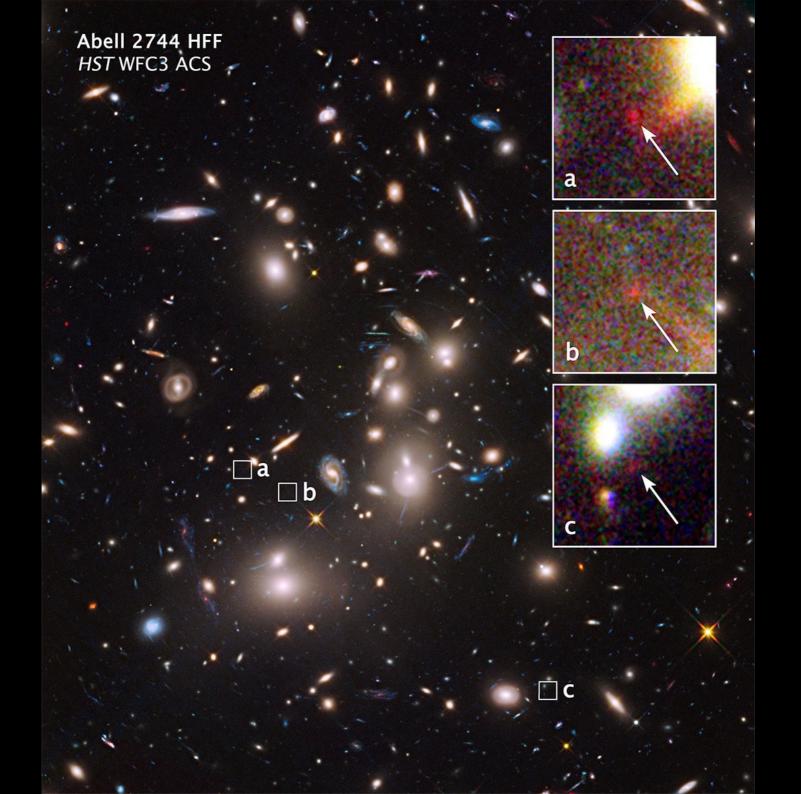


- Parallax- good for nearby stars
- Main sequence fitting- Good for nearby galaxies where individual stars can be resolved
- Variable stars: Cepheids- supergiant variable stars (Hubble used these).



Supernova- Type I: exploding White Dwarfs





- Parallax- good for nearby stars
- Main sequence fitting- Good for nearby galaxies where individual stars can be resolved
- Variable stars: Cepheids- supergiant variable stars
- Supernova- Type I: exploding White Dwarfs
- Redshift (we will get to this)

- Parallax- 1,000 pc (soon 1 Mpc with Gaia)
- Main sequence fitting- 2-3 Mpc
- Variable stars: ~50 Mpc
- Supernova- Type I: Any
- Redshift: Any

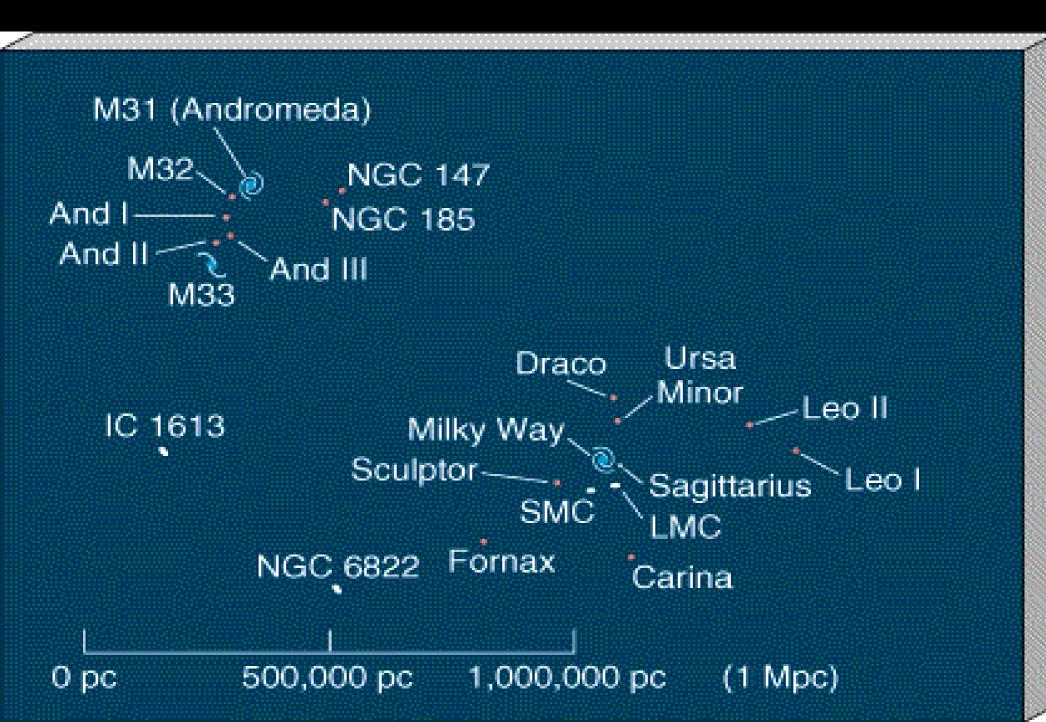
Cosmology.

Cosmology is the study of the Universe as a whole. It is about objects larger than individual galaxies and the evolution of the Universe and its contents.

The Local Group:

Our galaxy is a member of a small group (38) of gravitationally bound galaxies. There are 3 big spiral galaxies: The Milky Way, Andromeda, and Triangulum. Each of these 3 big galaxies have many smaller, irregular or dwarf elliptical galaxies orbiting them.

The Local Group:



Milky Way NGC NGC 185 147

Andromeda

NGC



