## - Previously



## Edwin Hubble's Classification Scheme



Sc
Sb
Ellipticals
Sa


Spirals


SBc

## Irregular galaxies:

These are not on the Hubble tuning fork. They do not have a disk, nor are they elliptical.
They tend to be young, gasrich and contain Pop I stars.

## Distances. How far away is everything and how can I tell?

## We use something called the distance ladder.

That is, each step depends on the previous one.


## Distances.

## How faf away is everything and how can I tell?

There will be summary slides

## Extra-Galactic Distances

- Parallax- good for nearby stars. GAIA- half our galaxy.



## Extra-Galactic Distances

- Parallax- good for nearby stars
- Main sequence fitting- Good for nearby galaxies where individual stars can be resolved (true for eclipsing binaries too).



## Extra-Galactic Distances

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- Variable stars: Cepheids- supergiant variable stars (Hubble used these).




## Extra-Galactic Distances

## Supernova- Type I: exploding White Dwarfs

Any distance


## Extra-Galactic Distances

- Parallax- good for nearby stars
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- Supernova- Type I: exploding White Dwarfs
- Redshift (we will get to this)


## Extra-Galactic Distances

- Parallax- $1,000 \mathrm{pc}$ (soon 1 Mpc with Gaia)
- Main sequence fitting- 2-3 Mpc
- Variable stars: $\sim 50 \mathrm{Mpc}$
- Supernova- Type I: Any
- Redshift: Any


## Cosmology.

Cosmology is the study of the Universe as a whole. It is about objects larger than individual galaxies and the evolution of the Universe and its contents.

## The Local Group:

Our galaxy is a member of a small group (38) of gravitationally bound galaxies. There are 3 big spiral galaxies: The Milky Way, Andromeda, and Triangulum.
Each of these 3 big galaxies have many smaller, irregular or dwarf elliptical galaxies orbiting them.

NGC 6822 Fornax $\quad$ Carina
0 pc
$500,000 \mathrm{pc}$
$1,000,000$ pc $\quad$ (1 Mpc)


## Triangulum



